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## Ph 203 (L4)

From last time:  
"Most" general  $V_{NN}$  consistent w symmetries & invariance:

$$V_{NN} = V_0(r) + V_1(r) (\hat{\sigma}_1 \cdot \hat{\sigma}_2) + V_2(r) (\hat{\tau}_1 \cdot \hat{\tau}_2) + V_3(r) (\hat{\sigma}_1 \cdot \hat{\sigma}_2) (\hat{\tau}_1 \cdot \hat{\tau}_2)$$

Central Force

$$+ V_T(r) \hat{S}_{12} + V_{T2}(r) \hat{S}_{12} (\hat{\tau}_1 \cdot \hat{\tau}_2)$$

Tensor Force

$$+ V_{LS}(r) \hat{L} \cdot \hat{S} + \dots \text{quadratic LS, } \hat{L} \cdot \hat{S} (\hat{\tau}_1 \cdot \hat{\tau}_2), (\hat{\sigma}_1 \cdot \hat{p})(\hat{\sigma}_2 \cdot \hat{p}), \dots$$

Velocity dependent Force

Can we build nuclei using  $V_{NN}$ ?  $\Rightarrow$  Maybe  
Why do we care?

### Astrophysics

1. Neutron Star properties/structure

$\rightarrow$  depends on Nuclear Eq. of State (NEOS)

$\rightarrow$  Nuclear matter properties vs  $\rho_n, \rho_p$

from subnuclear to supranuclear

"pasta"

$\rho > \rho_{Pb}$

2. Production of Elemental Abundances

$\rightarrow$  Nucleosynthesis  $\rightarrow$  Why more O than F?

$\rightarrow$  Big Bang Nucleosynthesis (BBN)

Fe " Co ?

$\rightarrow$  depends on Masses, Lifetimes, ...

Pb " Bi ?

3. Supernovae (SN)

$\rightarrow$  Why do they explode?

#### 4. Neutrinos ← Big Bang

→ Solar  $\nu$ , SN  $\nu$ , BB  $\nu$

### Physics Beyond Standard Model

#### 1. Electroweak Interactions

→ CKM matrix: Nuclear  $\beta$ -decay gives  $V_{ud}$   
↳ "Nuclear corrections"

→ CP Violation: Electric Dipole Moments (EDM)  
↳ "Nucl. Corr." for  $n, p$ , nuclei to get  
↳ quark EDM

→ New particles ( $M > \text{TeV}$ ) modify nuclear matrix elements in precision measurements

#### 2. Neutrino Properties

→ Oscillation of  $\nu$  flavors

↳ need cross sections  $\sigma_{\nu+A \rightarrow X}$  where  
 $A = N + Z$  of nuclear target

→ Is  $\nu = \bar{\nu}$  (Majorana neutrino) violates  
Lepton Number Conservation  $\phi_{\nu}$  Double  $\beta$  decay  
↳  $m_{\nu}$  depends on "Nucl. corr."

#### 3. Dark Matter detection

→ Nuclear targets ~~need~~  $\sigma_{DM+A}$

### Solving <sup>Quantum</sup> the Many-Body Problem

↓ Techniques for Nuclear Force / Reactions used  
in Condensed Matter & AMO many-body problem

Need:

# Approaches to Nuclear Force (How to calculate nuclear W.F.F.)

We will discuss:

- ① Phenomenological  $V(\vec{r})$
- ② Boson Exchange Potentials (OBEP) one Boson exch. potential
- ③ Effective Field Theory (EFT)  $\Rightarrow$  use QCD

## ① Phenom. Potential

Start with  $V_{NN}$  above based on symmetries/Invariance (rotations, parity, Time Reversal, Isospin, ...)

$\rightarrow$  Must add 3 nucleon force (3NF) + maybe 4NF, ...

$\rightarrow$  Most successful is: Argonne V18 + 3NF

18 params      22 params  
 $\leftarrow$  previous  $V_{NN}$  (see page 1)

$\hookrightarrow$  Fit the params to existing data & predict...

$\hookrightarrow$  works well for N-N + few-nucleon properties

$\rightarrow$  How far can we go?  $\Rightarrow$  heaviest stable nucleus is

$\rightarrow$  In practice  $A \approx 20$  is possible. Why?

e.g.  $^{16}\text{O}$  w 8 basis states

$\hookrightarrow$  4 energy states + 2 spin states

most evaluate

$$N_{\text{basis}}^A = 8^{16} \approx 3 \cdot 10^{14}$$

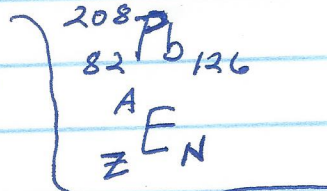
but  $^{76}\text{Ge}$  w 20 basis states

must consider  $20^{76} \approx 10^{99}$  (Ouch!)

$\hookrightarrow$  too much for today's computers

$\hookrightarrow$  maybe Quantum Computer?! see CANVAS

notes without  
3NF Binding  
Em of  $^3\text{He}$ ,  $^3\text{H}$   
off by  $\sim 1\text{MeV}$

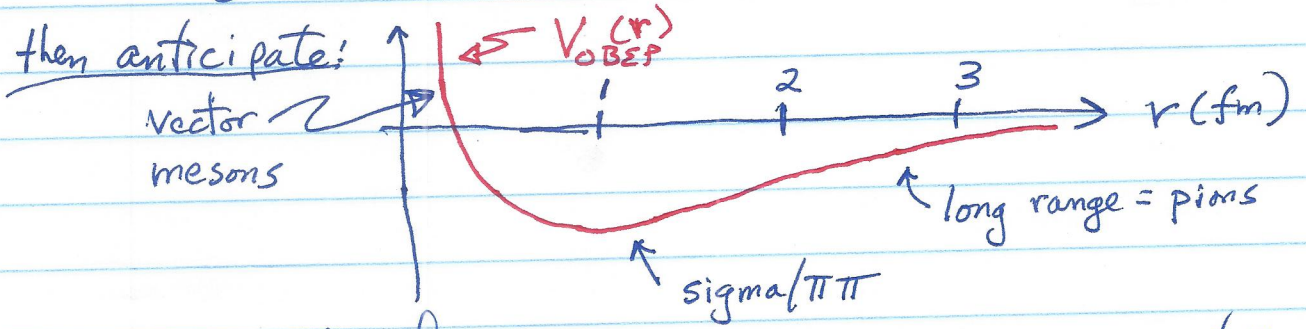


② OBEP Potential

↳ Yukawa-like force from massive particle exchange  
↳ gives short range  $V(r)$

Consider "light" mesons  $M \lesssim 1 \text{ GeV}$

| Meson                | $M$ (MeV)                      | $I J^{\pi}$ |              |
|----------------------|--------------------------------|-------------|--------------|
| $\pi$                | 138                            | $1 0^{-}$   | Pseudoscalar |
| $\eta$               | 548                            | $0 0^{-}$   |              |
| $a_0$                | 980                            | $1 0^{+}$   | Scalar       |
| $f_0/\sigma/\pi-\pi$ | 600<br>↳ but $\Gamma \sim 400$ | $0 0^{+}$   |              |
| $\rho$               | 770                            | $1 1^{-}$   | Vector       |
| $\omega$             | 782                            | $0 1^{-}$   |              |



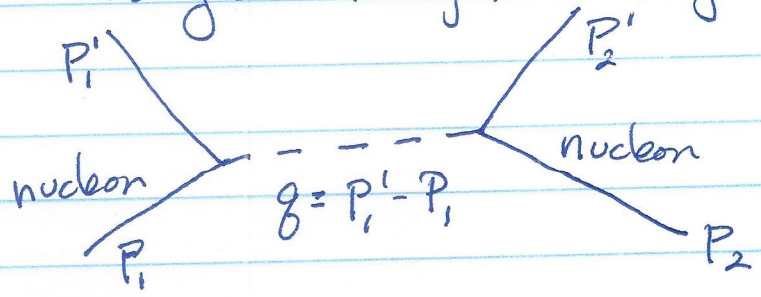
OBEP field theory

$$V_{NN}^{OBEP} = \sum_{i=\pi, \rho, \sigma, \dots} V_i(q)$$

momentum space

Use leading order Feynman diagrams:

@ tree-level



see App. D  
in Bert.

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Long Range:

then for pion exchange <sup>current</sup>

pure pseudoscalar gives wrong  $a_{\pi NN}$ , but  
pseudovector works  $\hat{=}$  gives  $\hat{S}_{12}$ !

$$\mathcal{L}_{\pi NN} = -\frac{f_{\pi NN}}{m_{\pi}} \bar{\psi} \gamma^{\mu} \gamma_5 \hat{\tau} \psi \cdot \partial_{\mu} \vec{\Phi}^{\pi}$$

which yields

$$V_{\pi NN}(q) \approx -\frac{f_{\pi NN}^2}{3m_{\pi}^2} \left( \frac{\vec{q}^2}{q^2 + m_{\pi}^2} \right) \left[ \hat{S}_{12}(q) - \hat{\sigma}_1 \cdot \hat{\sigma}_2 \right] \times \hat{\tau}_1 \cdot \hat{\tau}_2$$

gives via F.T. Fourier transform  $\frac{e^{-m_{\pi}cr/\hbar}}{r}$  ;  $\hat{S}_{12}(\vec{q}) = \frac{3}{q^2} (\hat{\sigma}_1 \cdot \vec{q})(\hat{\sigma}_2 \cdot \vec{q}) - \hat{\sigma}_1 \cdot \hat{\sigma}_2$   
 $\hat{=}$  FT of  $\hat{S}_{12}(q) \Rightarrow \hat{S}_{12}(r)$

Medium Range  $\Rightarrow$   $\sigma$ -meson:

$$\mathcal{L}_{\sigma NN} = -g_{\sigma} \bar{\psi} \psi \phi^{\sigma}$$

which yields:

$$V_{\sigma NN}(q) \approx \frac{g_{\sigma}^2}{q^2 + m_{\sigma}^2} \left( -1 - \frac{q^2}{2M_N^2} - \frac{\vec{L} \cdot \vec{S}}{2M_N^2} \right)$$

$\uparrow$  spin-orbit

Short Range :  $\Rightarrow$   $\omega$  meson

$$\mathcal{L}_{\omega NN} = -g_{\omega} \bar{\psi} \gamma^{\mu} \psi$$

which yields:

$$V_{\omega NN}(q) \approx \frac{g_{\omega}^2}{q^2 + m_{\omega}^2} \left( 1 - \frac{3\vec{L} \cdot \vec{S}}{M_N^2} \right)$$

$\uparrow$  repulsive core

lots of Params ( $g_i$ 's) to fit to lots of data  
but what are "theory" uncertainties??

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But! OBEP  $\neq$   $V_{NN}$

1. No handles for "theory" uncertainties

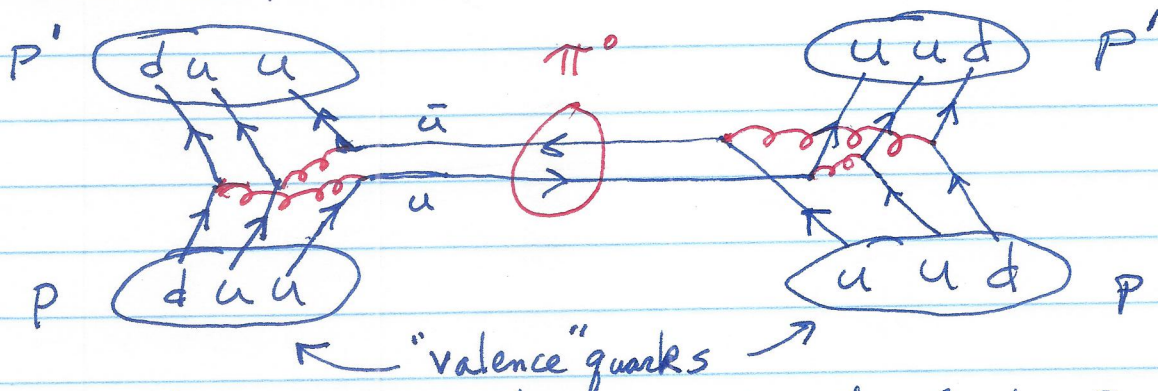
2. OBEP is a little ad hoc

↳ no real  $\sigma$ -meson

3. No QCD! the theory of strong int.

### ③ N-N in QCD

Consider  $p+p \rightarrow p+p$



Clearly this is tree-level (1-gluon), then for low E  $p+p$   
@ few MeV  $\alpha_s \geq 1$  (see later)  
↳ strong coupling constant

∴ many other diagrams needed

↳ Not perturbation theory

What to do?

1. Quark-based "Models" using  $q/g$  degrees of freedom  
not a real theory ← "naive"  $q$ -models - gluon

2. Try non-perturbative calcs. via "Lattice Gauge Theory"  
↳ only works "barely" for  $N+N$  → more later

3. Use symmetries (or lack thereof) in QCD to develop  
an Effective Field Theory (EFT)

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S. Weinberg ('91) said ...

*"If one writes down the most general possible Lagrangian, including all terms consistent with assumed symmetry principles, and then calculates matrix elements with this Lagrangian to any given order of perturbation theory, the result will simply be the most general possible S-Matrix consistent with analyticity, perturbative unitarity, cluster decomposition and assumed symmetry principles." - S. Weinberg*