

Ph 203 (L4)

From last time:

"Most" general V_{NN} consistent w symmetries & invariance:

$$V_{NN} = V_0(r) + V_\sigma(r) (\hat{\sigma}_1 \cdot \hat{\sigma}_2) + V_\tau(r) (\hat{\tau}_1 \cdot \hat{\tau}_2) + V_{\sigma\tau}(r) (\hat{\sigma}_1 \cdot \hat{\sigma}_2) (\hat{\tau}_1 \cdot \hat{\tau}_2)$$

Central Force

$$+ V_T(r) \hat{S}_{12} + V_{T\tau}(r) \hat{S}_{12} (\hat{\tau}_1 \cdot \hat{\tau}_2)$$

Tensor Force

$$+ V_{LS}(r) \hat{L} \cdot \hat{S} + \dots \text{quadratic LS, } \hat{L} \cdot \hat{S} (\hat{\tau}_1 \cdot \hat{\tau}_2), (\hat{\sigma}_1 \cdot \hat{p})(\hat{\sigma}_2 \cdot \hat{p}), \dots$$

Velocity dependent Force

Can we build nuclei using V_{NN} ? \Rightarrow Maybe
Why do we care?

Astrophysics

1. Neutron Star properties/structure

\rightarrow depends on Nuclear Eq. of State (NEOS)

\rightarrow Nuclear matter properties vs ρ_n, ρ_p

from subnuclear to supranuclear
"pasta" $\rho > \rho_{Pb}$

2. Production of Elemental Abundances

\rightarrow Nucleosynthesis \rightarrow Why more O than F?

\rightarrow Big Bang Nucleosynthesis (BBN)

\rightarrow depends on Masses, Lifetimes, ...

Fe " Co ?

Pb " Bi ?

3. Supernova (SN)

\rightarrow Why do they explode?

4. Neutrinos

← Big Bang

→ Solar ν , SN ν , BB ν

Physics Beyond Standard Model

1. Electroweak Interactions

→ CKM matrix: Nuclear β -decay gives V_{ud}
↳ "Nuclear corrections"

→ CP Violation: Electric Dipole Moments (EDM)
↳ "Nucl. corr." for n, p , nuclei to get
↳ quark EDM

→ New particles ($M > \text{TeV}$) modify nuclear matrix elements in precision measurements

2. Neutrino Properties

→ Oscillation of ν flavors

↳ need cross sections $\sigma_{\nu+A \rightarrow X}$ where

$A = N + Z$ of nuclear target

→ Is $\nu = \bar{\nu}$ (Majorana neutrino) violates
lepton Number Conservation $\phi \nu$ Double β decay

↳ m_ν depends on "Nucl. corr."

3. Dark Matter detection

→ Nuclear targets need σ_{DM+A}

Solving ^{Quantum} the Many-Body Problem

↳ Techniques for Nuclear Force / Reactions used
in Condensed Matter & AMO many-body problems

Need:

Approaches to Nuclear Force (How to calculate nuclear W.F.)

We will discuss:

- ① Phenomenological $V(\vec{r})$
- ② Boson Exchange Potentials (OBEP) one Boson
exch. potential
- ③ Effective Field Theory (EFT) \Rightarrow use QCD

① Phenom. Potential

Start with V_{NN} above based on symmetries/Invariance (rotations, parity, Time Reversal, Isospin, ...)

\rightarrow Must add 3 nucleon force (3NF) + maybe 4NF, ...

\rightarrow Most successful is: Argonne V18 + 3NF
18 params 22 params
 \leftarrow previous V_{NN} (see page 1)

\hookrightarrow Fit the params to existing data & predict ...

\hookrightarrow works well for N-N + few-nucleon properties

\rightarrow How far can we go? \Rightarrow heaviest stable nucleus is

\rightarrow In practice $A \approx 20$ is possible. Why?

e.g. ^{16}O w 8 basis states

\hookrightarrow 4 energy states + 2 spin states

most evaluate:

$$N_{\text{basis}}^A = 8^{16} \approx 3 \cdot 10^{14}$$

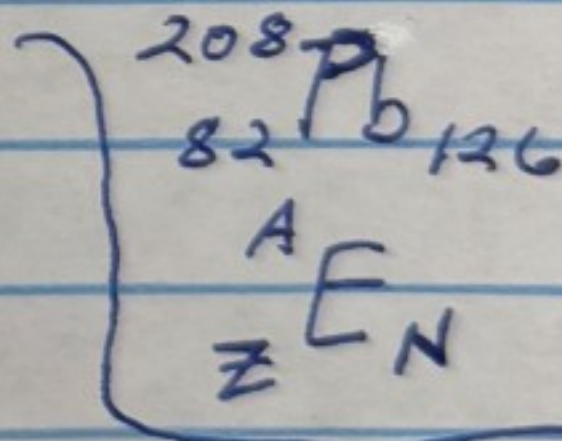
but ^{76}Ge w 20 basis states

must consider $20^{76} \approx 10^{99}$ (ouch!)

\hookrightarrow too much for today's computers

\hookrightarrow maybe Quantum Computer?! see CANVAS

note: without 3NF Binding Energy of ^3He , ^3H off by $\sim 1\text{MeV}$



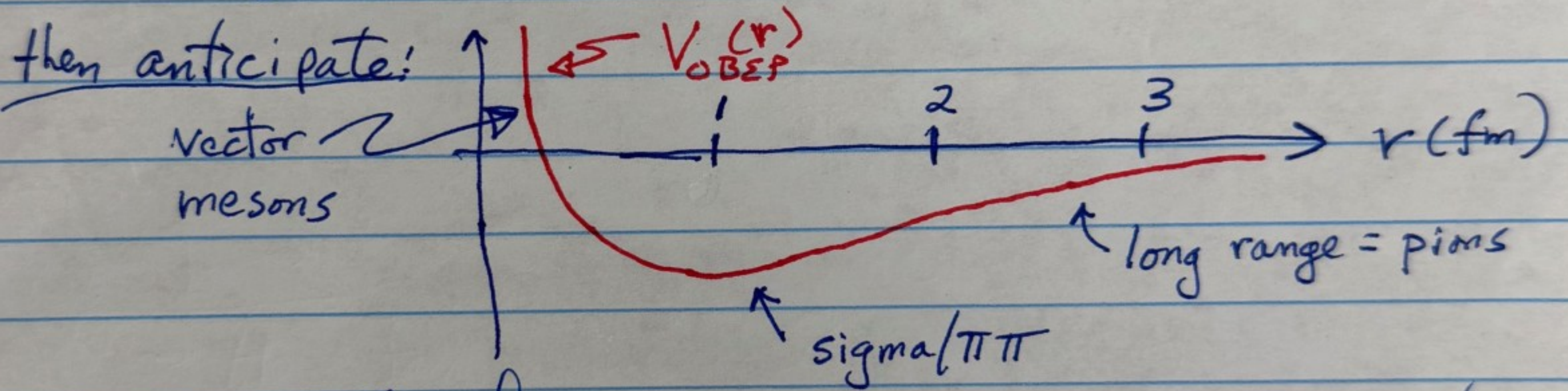
②

OBEP ^{ne meson exchange potential}

↳ Yukawa-like force from massive particle exchange
 ↳ gives short range $V(\vec{r})$

Consider "light" mesons $M \lesssim 1 \text{ GeV}$

Meson	$M \text{ (MeV)}$	$I J^\pi$	
π	138	$1 0^-$	} Pseudoscalar
η	548	$0 0^-$	
a_0	980	$1 0^+$	} Scalar
$f_0/\sigma/\pi-\pi$	600 ↳ but $\Gamma \sim 400$	$0 0^+$	
ρ	770	$1 1^-$	} Vector
ω	782	$0 1^-$	



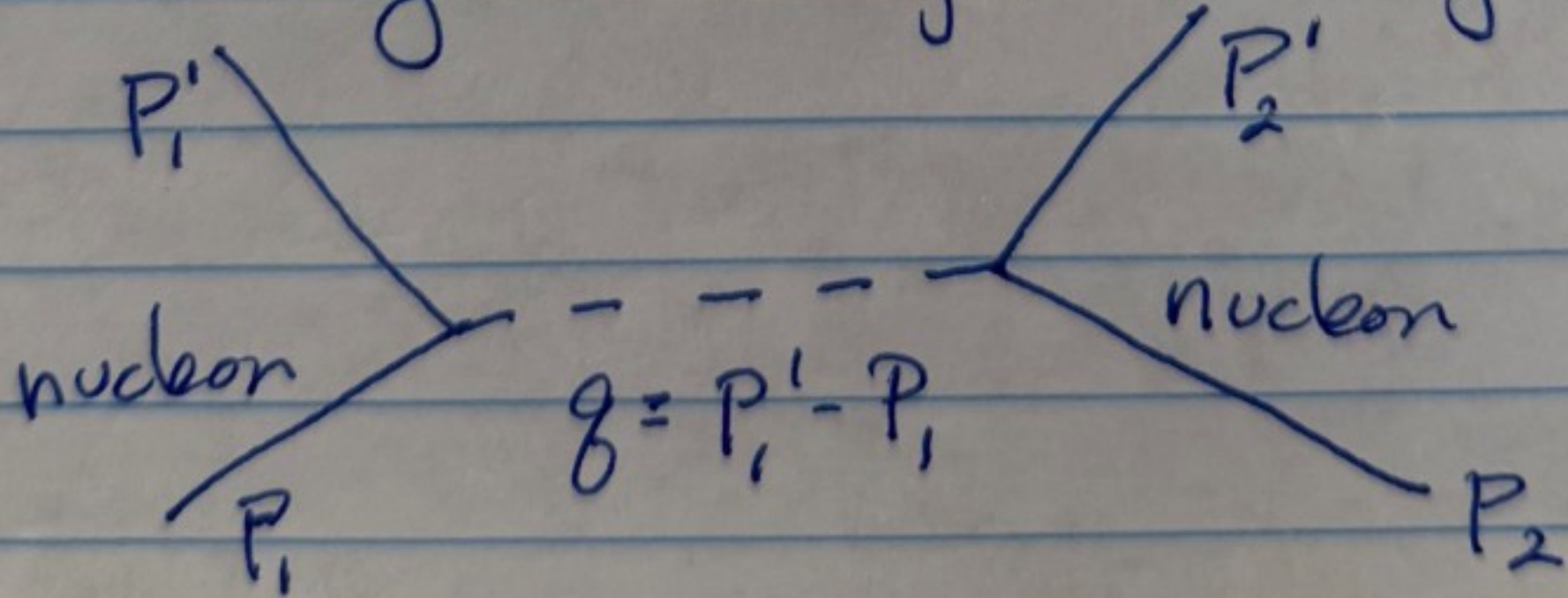
OBEP field theory

$$V_{NN}^{OBEP} = \sum_{i=\pi, \rho, \sigma, \dots} V_i(q)$$

momentum space

Use leading order Feynman diagrams:

@ tree-level



see App. D
in Bert.

Long Range:

then for pion exchange ^{current}
pure pseudoscalar gives wrong $a_{\pi N}$, but
pseudovector works $\hat{=}$ gives \hat{S}_{12} !

$$\mathcal{L}_{\pi NN} = -\frac{f_{\pi NN}}{m_{\pi}} \bar{\psi} \gamma^{\mu} \gamma_5 \hat{\partial} \psi \cdot \partial_{\mu} \vec{\Phi}^{\pi}$$

which yields

$$V_{\pi NN}(q) \hat{=} -\frac{f_{\pi NN}^2}{3m_{\pi}^2} \left(\frac{\vec{q}^2}{q^2 + m_{\pi}^2} \right) \left[\hat{S}_{12}(q) - \hat{\sigma}_1 \cdot \hat{\sigma}_2 \right] \hat{\tau}_1 \cdot \hat{\tau}_2$$

gives
via F.T.
Fourier transform

$$\frac{e^{-\frac{m_{\pi} c r}{\hbar}}}{r}$$

$$\hat{S}_{12}(q) = \frac{3}{q^2} (\hat{\sigma}_1 \cdot \vec{q})(\hat{\sigma}_2 \cdot \vec{q}) - \hat{\sigma}_1 \cdot \hat{\sigma}_2$$

$\hat{=}$ FT of $\hat{S}_{12}(q) \Rightarrow \hat{S}_{12}(r)$

Medium Range \Rightarrow σ -meson:

$$\mathcal{L}_{\sigma NN} = -g_{\sigma} \bar{\psi} \psi \phi^{\sigma}$$

which yields:

$$V_{\sigma NN}(q) \hat{=} \frac{g_{\sigma}^2}{q^2 + m_{\sigma}^2} \left(-1 - \frac{q^2}{2M_N^2} - \frac{\vec{L} \cdot \vec{S}}{2M_N^2} \right)$$

\uparrow spin-orbit

Short Range \Rightarrow ω meson

$$\mathcal{L}_{\omega NN} = -g_{\omega} \bar{\psi} \gamma^{\mu} \psi$$

which yields:

$$V_{\omega NN}(q) \hat{=} \frac{g_{\omega}^2}{q^2 + m_{\omega}^2} \left(1 - \frac{3\vec{L} \cdot \vec{S}}{M_N^2} \right)$$

\uparrow repulsive core

lots of params (g_i 's) to fit to lots of data
but what are "theory" uncertainties??

But! OBEP & V_{NN} have Problems!

1. No handles for "theory" uncertainties

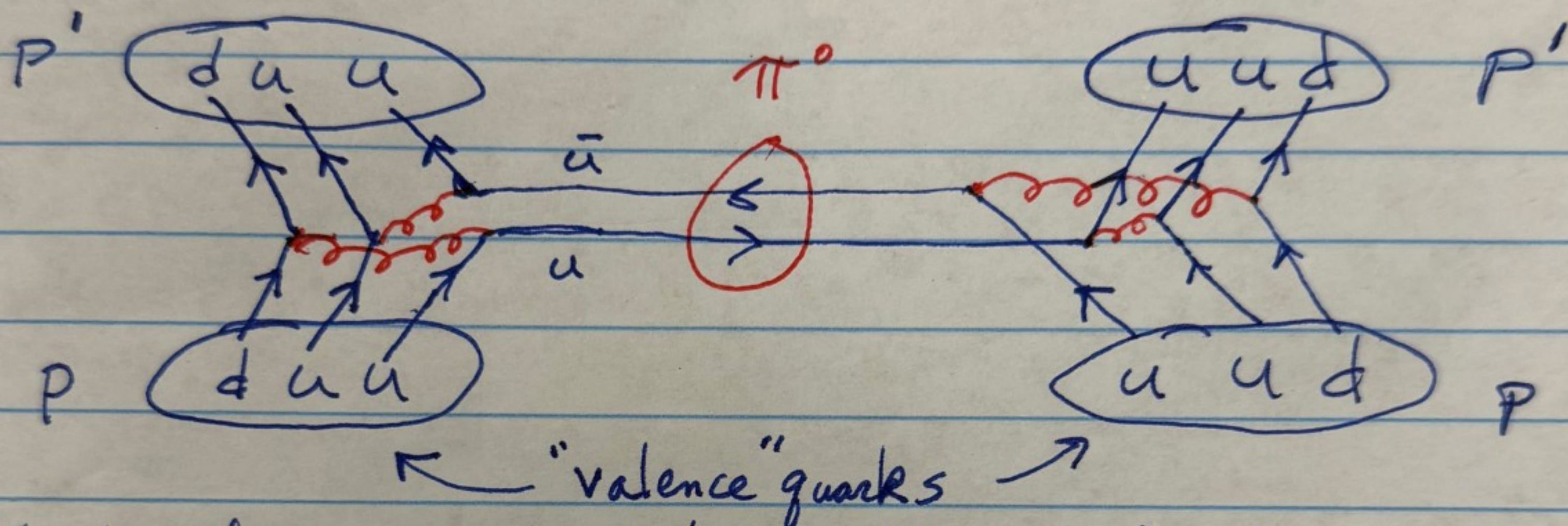
2. OBEP is a little ad hoc

↳ no real σ -meson

3. No QCD! the theory of strong int.

③ N-N in QCD

Consider $p+p \rightarrow p+p$



Clearly this is tree-level (1-gluon), then for low E $p+p$
@ few MeV $\alpha_s \geq 1$ (see later)
↳ strong coupling constant

∴ many other diagrams needed

↳ Not perturbation theory

What to do?

1. Quark-based "Models" using q/g degrees of freedom
not a real theory ← "naive" q -models - gluon

2. Try non-perturbative calcs. via "Lattice Gauge Theory"
↳ only works "barely" for $N+N$ → more later

3. Use symmetries (or lack thereof) in QCD to develop
an Effective Field Theory (EFT)

S. Weinberg ('91) said ...

"If one writes down the most general possible Lagrangian, including all terms consistent with assumed symmetry principles, and then calculates matrix elements with this Lagrangian to any given order of perturbation theory, the result will simply be the most general possible S-Matrix consistent with analyticity, perturbative unitarity, cluster decomposition and assumed symmetry principles." - S. Weinberg